

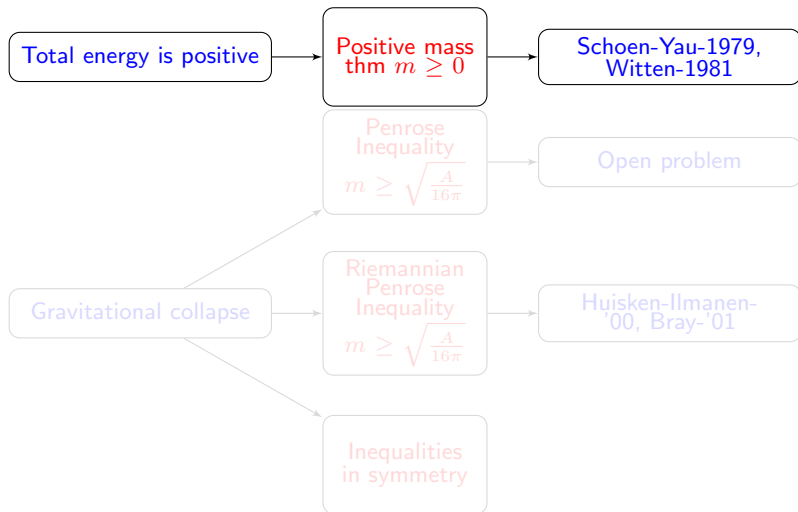
Mass functional and mass-angular momenta inequality for $U(1)^2$ -invariant black holes

Aghil Alae

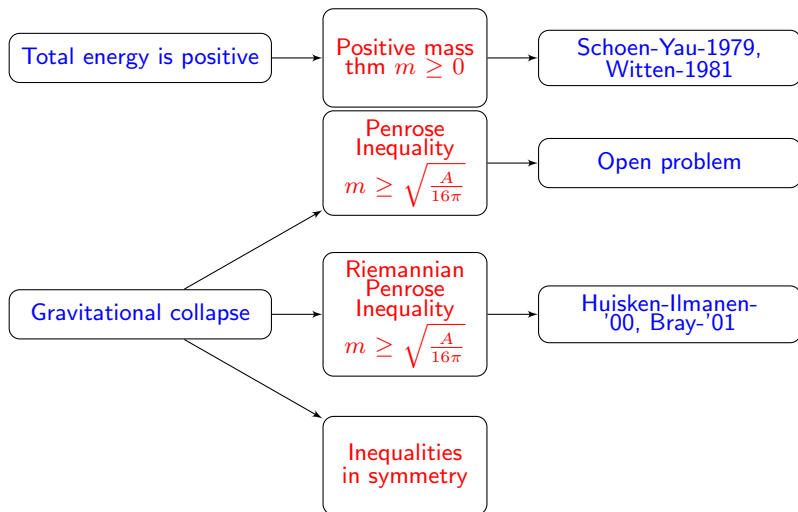
Department of Mathematics and Statistics
Memorial University of Newfoundland

Brandenburg an der Havel, Germany
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(joint work with Dr. Hari K. Kunduri)

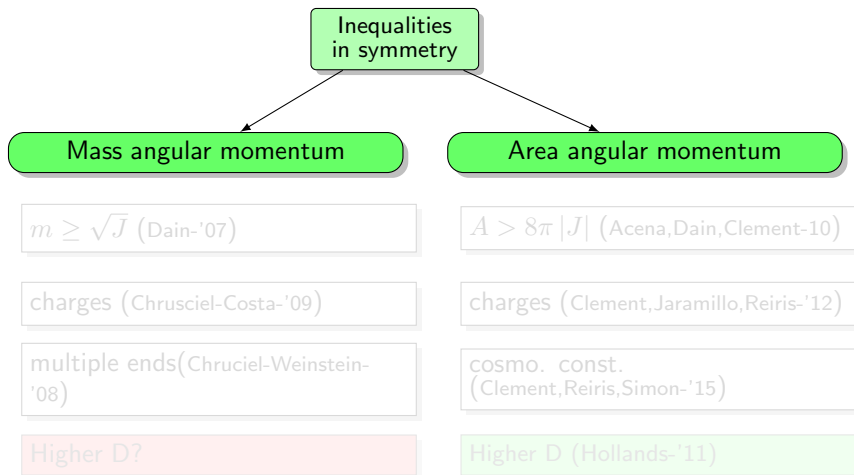
Motivation: Geometric Inequalities



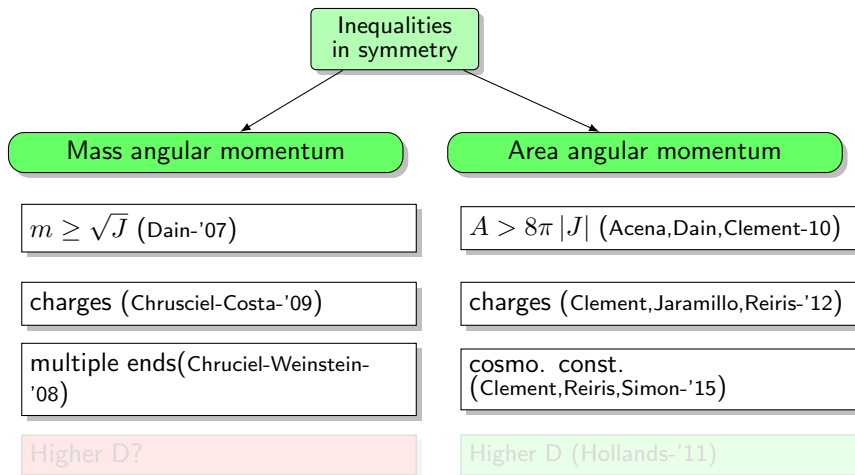
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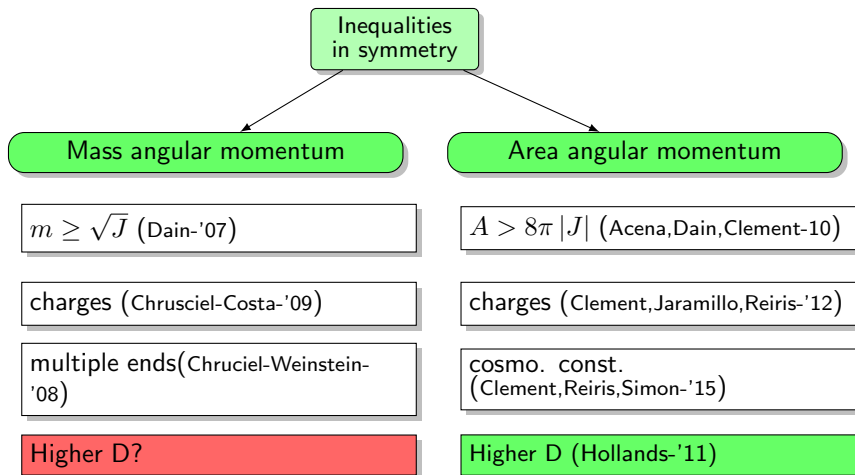
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Outline

- ▶ Mass and symmetry
- ▶ Review 4D mass-angular inequality
- ▶ Initial data for 5D BHs
- ▶ Main results
- ▶ Current and Future projects

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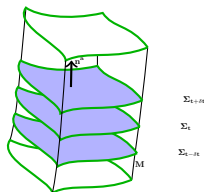
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Asymptotically flat and Constraint equation

- Assume (Σ, h) is smooth n -dimensional Riemannian manifold and C is a compact sub-manifold, (Σ, h) is **AF** if $\Sigma \setminus C$ is diffeomorphic to $\mathbb{R}^n \setminus B^n(0)$ for large r + **fall-offs**
- Consider spacetime (M, g) foliated by spacelike leaves (slices) (Σ, h_{ab}, K_{ab}) , where h_{ab} is induced metric on Σ and K_{ab} is extrinsic curvature tensor



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- Vacuum Einstein equations equivalent to evolution equations for initial data (Σ, h_{ab}, K_{ab}) + **2 constraint equations** on Σ

$$R_h - K^{ab}K_{ab} + \text{Tr}_h K = 0 \quad \text{Hamiltonian Constraint} \quad (1)$$

$$\nabla^b [K_{ab} - (\text{Tr}_h K)h_{ab}] = 0 \quad \text{Momentum Constraint} \quad (2)$$

Mass formulas

- In GR \implies mass has complicated quasi-local definition
 - Komar mass \implies AF and stationary sp-t at spacelike infinity

$$M_K = -\frac{1}{8\pi} \lim_{r \rightarrow \infty} \oint_{S_r} \star d\xi \quad \iff \quad \xi \text{ timelike KVF}$$

- ADM mass \implies AF sp-t and at spacelike infinity

$$M_{\text{ADM}} = \frac{1}{16\pi} \lim_{r \rightarrow \infty} \oint_{S_r} (\partial_c h_{ac} - \partial_c h_{aa}) n^c dS$$

- Bondi-Sachs mass \implies AF sp-t and at null infinity = ADM for stationary sp-t
- Hawking mass \implies AF sp-t and is not necessarily positive
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Symmetries on data

- An D dimensional initial data (Σ, h_{ab}, K_{ab}) is called **axisymmetric** if there exist $D - 2$ rotational KVFs ϕ^i which generates $U(1)^{D-2}$ isometry group on Riemannian manifold Σ and

$$\mathcal{L}_{\phi^i} h_{ab} = \mathcal{L}_{\phi^i} K_{ab} = 0$$

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Dain's mass functional

- **Dain ('05)** If $(\Sigma, h_{ab}, \bar{K}_{ab})$ is complete, $t - \phi$ symmetric, maximal, AF data in vacuum with two ends and rescaling

$$h_{ab} = e^{4v} \tilde{h}_{ab} \quad \tilde{h}_{ab} = e^{2q} (d\rho^2 + dz^2) + \rho^2 d\phi^2 \longleftarrow \boxed{\text{2 functions}}$$

then

- 1 \bar{K}_{ab} can be represented by **an scalar** Y
- 2 **positive definite** mass functional is

$$\mathcal{M}(v, Y) = \frac{1}{32\pi} \int_{\mathbb{R}^3} (16(dv)^2 + \rho^{-4} e^{-8v} (dY)^2) d\mu_0 \geq 0$$

- 3 \mathcal{M} evaluates ADM mass of these class
- 4 mass of any axisymmetric data $\geq \mathcal{M} \geq 0$
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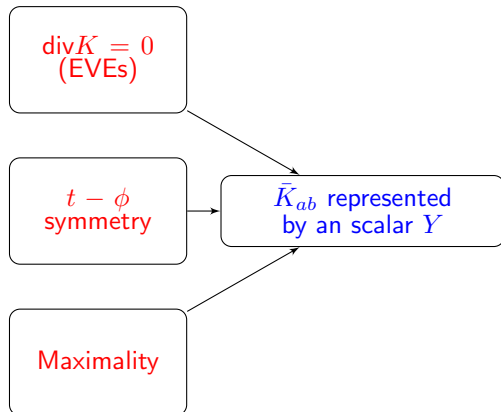
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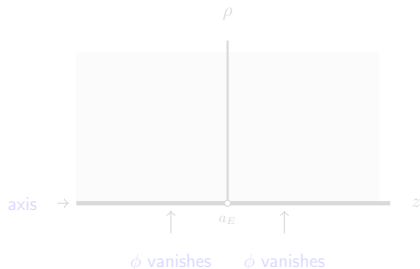
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Fundamental assumptions

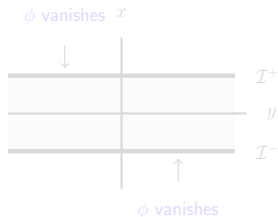


Orbit space of 3D initial data of BHs

- Σ with isometry group $U(1)$: $\frac{\Sigma}{U(1)} \cong \mathcal{B}$ orbit space which
- \mathcal{B} is, two dimensional manifold with boundary, homemorphic to upper-half plane.



(a) Orbit space as half plane

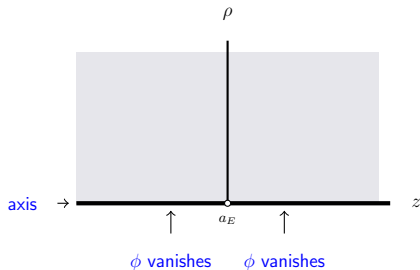


(b) Orbit space as infinite strip

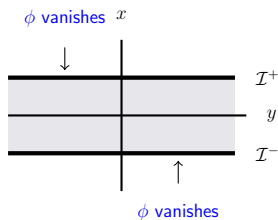
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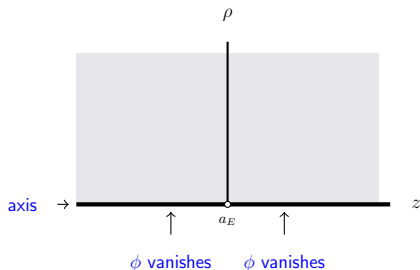


(d) Orbit space as infinite strip

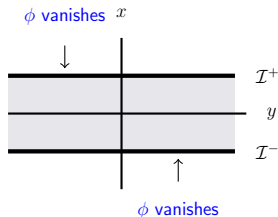
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(e) Orbit space as half plane



(f) Orbit space as infinite strip

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Results of Dain's mass functional

- Let (Σ, h, K) be a vacuum, AF, maximal, initial data set with appropriate decay. Then

$$m \geq \sqrt{J} = m_{\text{ex}} \quad (\text{Dain [local-'06,global-'07]}) \quad (3)$$

Moreover, the equality holds if and only if the data are a slice of [the extreme Kerr spacetime](#).

- (Dain '08) \mathcal{M} is a conserved quantity under axisymmetric evolution of Einstein equations.
- (Dain '14) Axisymmetric linear gravitational perturbations of the extreme Kerr black hole is stable.

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Restriction of method on dimension

- One of the restriction of Dain's method is existence of $U(1)^{D-2}$ -isometry on D -dimensional initial data (Why?)
- If we have an n -dimensional sp-t with $U(1)^{n-3}$ -isometry group \implies only for $n = 4, 5$ sp-t is AF
- We only consider $D = 4$ dimensional initial data (Σ, h, K)
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Topology of 5D black holes

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Horizon topology theorem (Galloway,Schoen-'06)

$H \cong S^3$ (and quotients), $H \cong S^1 \times S^2$ and connected sums of these two cases

Topology of Stationary Rotating BH (Hollands,Holland,Ishibashi-'10)

- 1 $H \cong \#k(S^1 \times S^2) \#l$ Lens space, $k \geq 0$, $l \geq 1$
- 2 $H \cong$ quotients of S^3

Topological censorship (Galloway, Friedmann, Schleich, Witt)

Domain of outer communication is simply connected

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Domain of outer communication (DOC)

- (Orlik, Raymond-'70) If Σ is a 4D simply connected manifold with action $U(1)^2$ then

$$\Sigma \cong \#n(S^2 \times S^2) \#n'(\pm\mathbb{C}P^2)$$

- (Hollands, Holland, Ishibashi, '12) If Σ is a 4D simply connected AF manifold with action $U(1)^2$ then

$$\Sigma \cong \mathbb{R}^4 \#n(S^2 \times S^2) \#n'(\pm\mathbb{C}P^2)$$

Domain of outer communication (DOC)

Theorem(Hollands,Holland,Ishibashi,'12): Consider stationary, rotating vacuum black hole. Then DOC has topology $DOC \cong \mathbb{R} \times \Sigma$ where

$$\Sigma \cong \mathbb{R}^4 \# n (S^2 \times S^2) \# n' (\pm \mathbb{C}P^2) \setminus B \quad (4)$$

where B is compact, $\partial B \cong H$. B is **'black hole region'**.

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Remark Alae, Kunduri, Martinez-Pedroza ('13)

Question Does topology of H uniquely determine Σ ?

- 1 Choice of H does not specify a unique smooth submanifold B in $\mathbb{R}^4 \implies$ different B can yield different Σ .
- 2 They are unique up to homological level.

Question Does B uniquely embedded in \mathbb{R}^4 ?

- 1 No, there are different possibilities.
- 2 But it is un-knotted

Remark Alae, Kunduri, Martinez-Pedroza ('13)

Question Does topology of H uniquely determine Σ ?

- 1 Choice of H does not specify a unique smooth submanifold B in $\mathbb{R}^4 \implies$ different B can yield different Σ .
- 2 They are unique up to homological level.

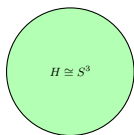
Question Does B uniquely embedded in \mathbb{R}^4 ?

- 1 No, there are different possibilities.
- 2 But it is un-knotted

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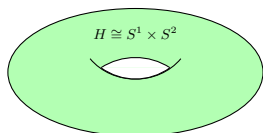
- If $H \cong S^3$ there is a unique possibility for black hole region $B \cong B^4$ (4-ball)
- If $H \cong S^1 \times S^2$ the **standard choice** $B \cong S^1 \times B^3$
- For multiple black holes one can find a standard choice!

Slice of 5D BH: Alae, Kunduri, Martinez-Pedroza ('13)



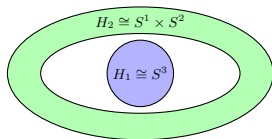
Myers-Perry BH

$$\Sigma \cong \mathbb{R} \times S^3, \chi = 0$$



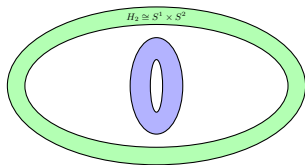
Black Ring

$$\Sigma \cong (S^2 \times D^2) \# \mathbb{R}, \chi = 1$$



Black Saturn

$$\Sigma \cong (S^2 \times D^2) \# \mathbb{R} \# B^4, \chi = 0$$

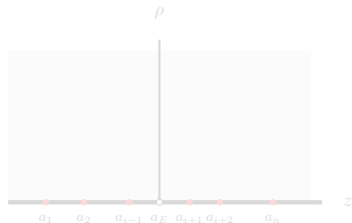


Bicycling Black Rings

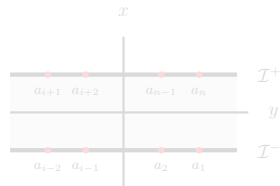
$$\Sigma \cong \#2(S^2 \times D^2) \# \mathbb{R}, \chi = 1$$

Topology of Σ and Orbit space of 5D BHs

- Σ with isometry group $U(1)^2$: $\frac{\Sigma}{U(1)^2} \cong \mathcal{B}$ orbit space
- \mathcal{B} is two dimensional manifold with boundary and corners (Hollands-Yazadjiev ('08) for sp-t).



(g) Orbit space as half plane



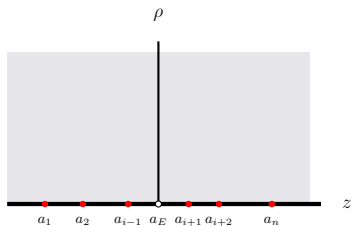
(h) Orbit space as infinite strip

- Orbit space structure of Σ is related to topology of H and bubbles

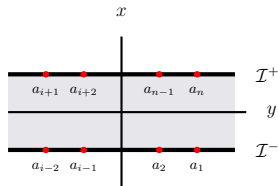
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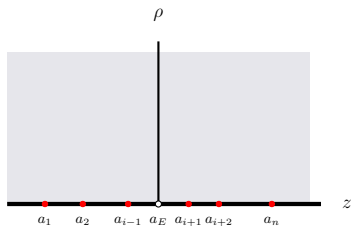
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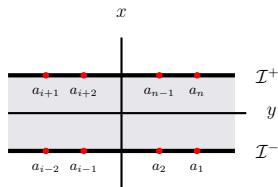
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Result

- Alaee-Kunduri ('14) If $(\Sigma, h_{ab}, \bar{K}_{ab})$ is complete, $t - \phi^i$ symmetric, maximal, AF data in vacuum with rescaling

$$h_{ab} = \Phi^2 \tilde{h}_{ab} \quad \tilde{h}_{ab} = e^{2U} (d\rho^2 + dz^2) + \lambda_{ij} d\phi^i d\phi^j \leftarrow \boxed{\text{4 functions}}$$

where $\rho^2 = \det \lambda_{ij}$. Then

- \bar{K}_{ab} is divergence-less and traceless 2 tensor and can be represented by two scalars Y^i and λ_{ij}

$$\bar{K}_{ab} \equiv P_{(a}^1 \phi_{b)}^1 + P_{(a}^2 \phi_{b)}^2$$

$$P \equiv \lambda^{-1} S \quad \text{where} \quad S_a^i \equiv \frac{1}{2\rho^2} i_{\phi^2} i_{\phi^1} \star dY^i, \quad d \star S^i = 0$$

- If (Σ, h_{ab}, K_{ab}) is a AF, $U(1)^2$ -invariant, maximal data in vacuum then

$$K_{ab} K^{ab} \geq \bar{K}_{ab} \bar{K}^{ab}$$

maximal, AF, $t - \phi^i$
symmetric initial data

6 functions
with a
constraint

Mass functional \mathcal{M}

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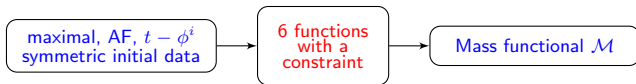
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4 dimensional mass functional

- **Alaee-Kunduri ('14)** If $(\Sigma, h_{ab}, \bar{K}_{ab})$ is complete, $t - \phi^i$ symmetric, maximal, AF data in vacuum, then
 - Mass functional is

$$\begin{aligned} \mathcal{M} &= \frac{\pi}{4} \int_{\mathcal{B}} \left(-\frac{\det d\lambda}{\rho^2} + e^{-6v} \frac{dY^t \lambda^{-1} dY}{2\rho^2} + 6 (dv)^2 \right) \rho d\rho dz \\ &\quad - \frac{\pi}{4} \sum_{\text{rods}} \int_{I_i} \log V_i dz \end{aligned}$$

where $v = \log \Phi$ and I_i are intervals with direction vector v^i on $\partial\mathcal{B}$

$$V_i = \frac{2\sqrt{\rho^2 + z^2 \lambda_{ij} v^i v^j}}{\rho^2} \quad \text{where } z \in (a_i, a_{i+1}) \text{ and } \rho \rightarrow 0$$

Results of \mathcal{M}

- 1 \mathcal{M} encodes information of rod structure
- 2 \mathcal{M} is a positive definite functional for a large class(Which?)
- 3 \mathcal{M} evaluates ADM mass of these class
- 4 mass of any axisymmetric data $\geq \mathcal{M}$
- 5 critical points of \mathcal{M} are stationary sp-t
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Question

- Geometric inequalities in vacuum for two known extreme AF solutions

$$m^3 \geq \frac{27\pi}{32} (|J_1| + |J_2|)^2 \quad (\text{Myers-Perry})$$

$$m^3 \geq \frac{27\pi}{4} |J_1| (|J_2| - |J_1|) \quad (\text{black ring})$$

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Uniqueness of extreme $5D$ BHs

Theorem [Figueras, Lucietti-'10] Consider a five dimensional, AF, stationary black hole solution of the vacuum Einstein equations, with $U(1)^2$ isometry and a connected degenerate horizon (with non-toroidal sections). There exists at most one such solution with given angular momenta J_1, J_2 and a given interval structure.

- The non-extreme version was proved by Hollands-Yazadjiev ('08)

Extreme class of data-Alaee-Kunduri 15

Definition: The set of *extreme class E* is the collection of data arising from extreme, vacuum, AF, $\mathbb{R} \times U(1)^2$ invariant black holes which consist of triples $u_0 = (v_0, \lambda'_0, Y_0)$ where v_0 is a scalar, $\lambda'_0 = [\lambda_{ij}]$ is a positive definite 2×2 symmetric matrix, and Y_0 is a column vector with the appropriate bounds.

Main result-Alaee-Kunduri '15

Theorem Let (Σ, h_{ab}, K_{ab}) be an AF, maximal, $U(1)^2$ -invariant, vacuum initial data with mass m and fixed angular momenta J_1 and J_2 and fixed orbit space \mathcal{B} . Then in small neighborhood

$$m \geq f(J_1, J_2)$$

for some f which depends on the orbit space \mathcal{B} . Moreover, $m = f(J_1, J_2)$ in the neighborhood if and only if the data are extreme data.

- Currently we are investigating the global mass-angular momenta for the large class of data (including MP data)

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Open Problems for 5D BHs

- Positivity of mass functional \mathcal{M} for all orbit spaces and extend the global proof to all orbit spaces.
- Mass-angular momenta-charge?
- Mass-angular momenta with multiple ends?
- Is \mathcal{M} a conserved quantity under $U(1)^2$ -invariant evolution of Einstein equations?
- Study stability (or instability) of $U(1)^2$ -invariant linear gravitational perturbations of the extreme MP and extreme doubly spinning black ring.

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Thank you

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