## Geometric invariants on black hole initial data

Robert Sansom

#### Queen Mary University of London

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Black hole initial data is initial data for the vacuum Einstein field equations that contains a marginally outer trapped surface (MOTS), also called an apparent horizon. Black hole initial data is initial data for the vacuum Einstein field equations that contains a marginally outer trapped surface (MOTS), also called an apparent horizon.

The MOTS is the boundary of the black hole in the initial data.



Black hole initial data is initial data for the vacuum Einstein field equations that contains a marginally outer trapped surface (MOTS), also called an apparent horizon.



For example, in the Schwarzschild spacetime the event horizon coincides with the apparent horizon.

The Schwarzschild spacetime is also stationary - i.e. there exists a time-like Killing vector field.

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Motivated by:

PRL 93, 231101 (2004)

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#### A New Geometric Invariant on Initial Data for the Einstein Equations

Sergio Dain Albert-Einstein-Institut, am Mühlenberg 1, D-14476, Golm, Germany (Received 14 July 2004; published 1 December 2004)

For a given asymptotically flat initial data set for Einstein equations a new geometric invariant is constructed. This invariant measures the departure of the data set from the stationary regime; it vanishes if and only if the data are stationary. In vacuum, it can be interpreted as a measure of the total amount of radiation contained in the data.

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## The approximate Killing equation

Given  $h_{ab} \in \mathcal{M}_2$  space of Riemannian metrics,  $K_{ab} \in \mathcal{S}_2$  space of symmetric 2-tensors.

Let  ${\mathscr C}$  and  ${\mathscr X}$  denote the spaces of scalars and vectors.

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Let  ${\mathscr C}$  and  ${\mathscr X}$  denote the spaces of scalars and vectors.

The constraint map is the map  $\Phi: \mathcal{M} \times \mathcal{S}_2 \to \mathcal{C} \times \mathfrak{X}$  with

$$\Phi\left(\begin{array}{c}h_{ij}\\K_{ij}\end{array}\right) \equiv \left(\begin{array}{c}r+K^2-K_{ij}K^{ij}\\-D^jK_{ij}+D_iK\end{array}\right)$$

Linearisation of the constraint map at  $(h_{ij}, K_{ij})$ :  $D\Phi: \mathscr{S}_2 \times \mathscr{S}_2 \to \mathscr{C} \times \mathscr{X}$ 

$$D\Phi\left(\begin{array}{c}\gamma_{ij}\\Q_{ij}\end{array}\right) = \left(\begin{array}{c}D^i D^j \gamma_{ij} - r_{ij} \gamma^{ij} - \Delta_h \gamma + H\\-D^j Q_{ij} + D_i Q - F_i\end{array}\right).$$

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The formal adjoint:

$$D\Phi^* \left(\begin{array}{c} X\\ X_i \end{array}\right) = \left(\begin{array}{c} D_i D_j X - Xr_{ij} - \Delta_h Xh_{ij} + H_{ij} \\ D_{(i}X_{j)} - D^k X_k h_{ij} + F_{ij} \end{array}\right).$$

where  $H, F_i, H_{ij}$  and  $F_{ij}$  are terms of lower order which vanish under time symmetry -  $K_{ij} = 0$ .

Surprisingly, the elements of the kernel of  $D\Phi^*$  are the symmetries of the spacetime determined by the initial data set (S, h, K). (Moncrief, 1975) Surprisingly, the elements of the kernel of  $D\Phi^*$  are the symmetries of the spacetime determined by the initial data set (S, h, K). (Moncrief, 1975)

In other words, a solution  $(N, N^i)$  to  $D\Phi^*(N, N^i) = 0$  also solves the Killing Initial data (KID) equations

$$NK_{ij} + D_{(i}N_{j)} = 0,$$
  
$$N^{k}D_{k}K_{ij} + D_{i}N^{k}K_{kj} + D_{j}N^{k}K_{ik} + D_{i}D_{j}N = N(r_{ij} + KK_{ij} - 2K_{ik}K^{k}_{j}).$$

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The spacetime evolving from the initial data will have a killing vector with lapse N and shift  $N^i.$  (Beig - Chrusciel, 1996)

We will make use of the following **Bartnik operators** related to the linearisation of the constraint map and its formal adjoint:

$$\mathcal{P}\left(\begin{array}{c}\gamma_{ij}\\q_{kij}\end{array}\right) \equiv D\Phi\left(\begin{array}{c}\gamma_{ij}\\-D^{k}q_{kij}\end{array}\right),$$
$$\mathcal{P}^{*}\left(\begin{array}{c}X\\X_{i}\end{array}\right) \equiv \left(\begin{array}{c}1&0\\0&D_{k}\end{array}\right) \cdot D\Phi^{*}\left(\begin{array}{c}X\\X_{i}\end{array}\right).$$

The approximate Killing operator  $: \mathscr{P} \circ \mathscr{P}^* : \mathscr{C} \times \mathfrak{X} \to \mathscr{C} \times \mathfrak{X}$ 

$$\mathcal{P} \circ \mathcal{P}^* \left( \begin{array}{c} X\\ X_i \end{array} \right) \equiv \left( \begin{array}{c} 2\Delta_h \Delta_h X - r^{ij} D_i D_j X + 2r \Delta_h X + \text{l.o.t} \\ \\ D^j \Delta_h D_{(i} X_{j)} + D_i \Delta_h D^k X_k + \text{l.o.t} \end{array} \right)$$

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Then the approximate Killing equation is

$$\mathscr{P} \circ \mathscr{P}^* \begin{pmatrix} X \\ X_i \end{pmatrix} = 0.$$
 (AKE)

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## Properties of (AKE)

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- A solution to the KID equations also solves (AKE),
- The approximate Killing operator is self-adjoint, fourth order and elliptic,
- (AKE) is the Euler-Lagrange equation of

$$\int_{S} \mathscr{P}^{*}(X, X_{i}) \cdot \mathscr{P}^{*}(X, X_{i}) \,\mathrm{d}\mu.$$

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Note: We cannot construct (AKE) using  $D\Phi^*$  because the functional

$$\int_{S} D\Phi^{*}\left(X, X_{i}\right) \cdot D\Phi^{*}\left(X, X_{i}\right) \mathrm{d}\mu$$

contains terms of inconsistent physical dimension.

## Solvability of (AKE)

- Focus on the symmetry corresponding to time translation,
- One asymptotic end,
- One (or several) inner bdrys that are marginally outer trapped surfaces (MOTS).



#### Assume that the initial data is asymptotically flat

$$h_{ab} - \delta_{ab} \in H^{\infty}_{-\frac{1}{2}} \qquad (= o_{\infty}(|x|^{-\frac{1}{2}}))$$
  
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$$K_{ab} \in H^{\infty}_{-\frac{3}{2}}$$

The initial data is called stationary if there exists  $(N,N^i)\in H^2_{1/2}$  such that  $\mathscr{P}^*(N,N^i)=0.$ 

#### Lemma

Assume on the bdry  $\partial S$  that one has

$$\begin{cases} N|_{\partial S} = 0\\ \Delta_h N|_{\partial S} = 0\\ N^i|_{\partial S} = 0\\ \not D N^i|_{\partial S} = 0 \end{cases}$$

then there exists a solution  $(N,N^i)\in H^\infty_{1/2}$  to (AKE) if and only if the data is stationary.

Where D is the covariant derivative along the normal direction to  $\partial S$ .

### Main existence theorem

### Theorem (S, Valiente Kroon '22)

Given smooth functions  $f,g,f^i,$  and  $h^i$  on  $\partial S,$  then the BVP

$$\begin{split} \mathcal{P} \circ \mathcal{P}^* \begin{pmatrix} X \\ X^i \end{pmatrix} &= 0 \qquad \text{on } S, \\ \begin{cases} X|_{\partial S} &= f, \\ \Delta_h X|_{\partial S} &= g, \\ X^i|_{\partial S} &= f^i, \\ \mathcal{D} X^i|_{\partial S} &= h^i, \end{split}$$

has a unique solution such that

$$\begin{split} X &= \lambda |x| + \vartheta, \qquad \vartheta \in H^\infty_{\frac{1}{2}} \\ X^i \in H^\infty_{\frac{1}{2}}. \end{split}$$

Futhermore,  $\lambda$  vanishes if the initial data is stationary.

The number  $\lambda$  is **Dain's invariant**.

• Under time symmetry,  $K_{ij} = 0$ , Dain's invariant can be written as the boundary integral

$$\lambda = -\frac{1}{8\pi} \oint_{S_{\infty}} n^a D_a \Delta_h X \mathrm{d}S,$$

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- It can be used to measure the deviation from stationarity at the asymptotic end.

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Futhermore,  $\lambda$  vanishes if the initial data is stationary.

## Specifying the boundary data

How do we turn the if in the previous theorem into an if and only if?

## 'How much' of the KID equations can be solved on $\partial S$ ?

To do this, perform a 2+1 decomposition of the KID equations.



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Now, assume time symmetry,  $K_{ij} = 0$ . This assumption sets  $X^i = 0$  on S.

Projecting the KID equations with the 2-metric:  $\bar{h}_{ij} = h_{ij} - \rho_i \rho_j$ onto  $\partial S$  obtains the intrinsic equation on  $\partial S$ :

$$\Delta_{\bar{h}}X - \frac{1}{2}(\bar{r} + |\bar{K}|^2)X = 0$$

where  $|\bar{K}|^2 = \bar{K}_{AB}\bar{K}^{AB}$ .

## Stability of MOTS



A MOTS evolves into a marginally outer trapped tube if it is **stable** - the lowest eigenvalue of the MOTS stability operator,  $\mathcal{L}$ , is positive

$$\mathcal{L} \equiv -\Delta_{\bar{h}} + \frac{1}{2}(\bar{r} - |\bar{K}|^2)$$

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KID equation on  $\partial S: \Delta_{\bar{h}} X - \frac{1}{2}(\bar{r} + |\bar{K}|^2) X = 0$ 

#### Lemma

If the MOTS is stable then the only solution to the intrinsic KID equation

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Setting X = 0 in the KID equations leaves only

$$D_i D_j X = 0$$

on  $\partial S.$  This suggests setting  $\Delta_h X=0$  as the boundary condition for (AKE).

The other components of the decomposition of the KID equations (normal-normal and normal-tangential etc.) leads to the following geometric invariant.

#### Lemma

Let  $X=\Delta_h X=0$  on  $\partial S.$  Then the KID equations are satisfied at  $\partial S$  if and only if  $\omega=0$  where

$$\omega = \oint_{\partial S} |\bar{K}|^2 |\mathcal{D}X|^2 \mathrm{d}S.$$

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$$\omega = \oint_{\partial S} |\bar{K}|^2 |\mathcal{D}X|^2 \mathrm{d}S.$$

Thus, in the time symmetric setting, The boundary conditions  $X|_{\partial S} = \Delta_h X|_{\partial S} = 0$  with  $\omega = 0$  is enough to ensure that  $\lambda = 0$  if and only if the data is stationary.

Conversely, if  $\omega \neq 0$  then the data is not stationary.

► The non-time symmetric case? Much harder to see how to specify boundary data because data must also be specified for  $X^i$ . For  $\vec{X} = (X, X^A)$  the intrinsic part of the KID equations can be written as an elliptic system

$$\Delta_{\bar{h}}\vec{X} + T \cdot \bar{D}\vec{X} + C \cdot \vec{X} = \vec{F}.$$

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The structure of this equation needs to be understood,

- Unique continuation of the KID equations from  $\partial S$  to S,
- Make connection with uniqueness of black holes using approximate symmetries,
- Evolution of Dain's invariant.

Thank you!